# The Radiation Environment in Space

## Sources of Radiation in Space

- Galactic Cosmic Rays
- The Sun
  - The Solar Wind
  - Coronal Mass Ejections

## Radiation in Space

NASA research programs for The International Space Station Mission to Mars

- Understand the biological effects
- Design adequate shielding

Source unknown

**Isotropic** 

Fluence constant over time

Figure removed due to copyright restrictions.

Figure 2.1 in [SSB-Crew Hazards].

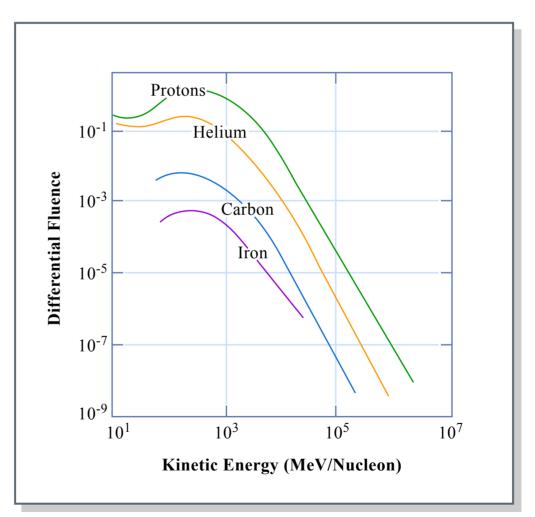
Commission on Physical Sciences, Mathematics, and Applications, Space Studies Board (SSB). *Radiation Hazards to Crews of Interplanetary Missions: Biological Issues and Research Strategies*. Washington, DC:

National Academies Press, 1996.

See http://books.nap.edu/books/0309056985/html/14.html#page bottom.

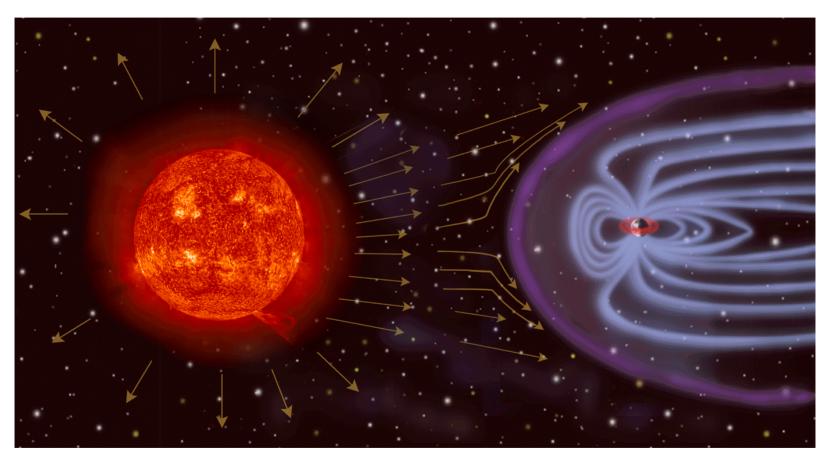
**Broad energy** distribution

Peak at about 1 GeV/nucleon



Attenuation in the upper atmosphere

Image removed due to copyright restrictions.



Courtesy of NASA. "Living in the Atmosphere of the Sun." [updated 20 Jan 2000, cited 29 March 2004.] http://www-istp.gsfc.nasa.gov/exhibit/main.html

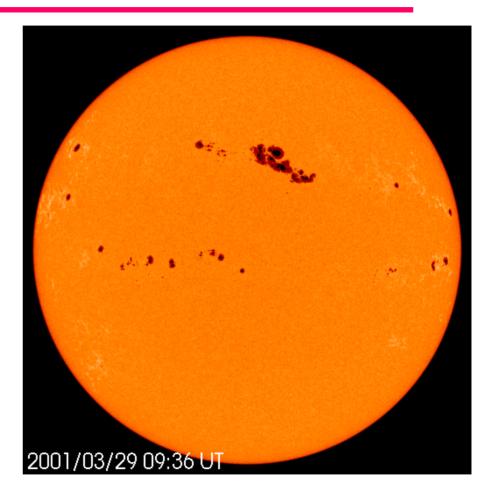
The solar wind deforms the Earth's magnetic field lines

Image removed due to copyright restrictions.

**Sunspots** 

**Observed for centuries** 

Occur in an 11 year cycle



Courtesy of NASA Goddard Space Flight Center. [updated 30 March 01, cited 29 March 2004.] http://www.gsfc.nasa.gov/gsfc/spacesci/solarexp/sunspot.htm

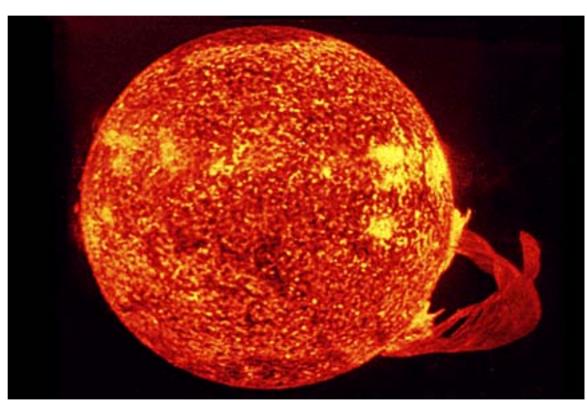
Image removed due to copyright restrictions.

During maximum sunspot activity, there is an increased likelihood of solar flares and coronal mass ejections.

#### Solar flares

Associated with production of high energy particles, mostly protons.

Potentially lethal doses



Courtesy of NASA. "Our Magnificent Sun." [cited 29 March 2004] http://cossc.gsfc.nasa.gov/images/epo/gallery/solar/
1996 photo from Skylab

## Violent Space Weather

**High Energy Coronal Mass Ejections produce "shocks"** 

- Highest energy particles reach Earth in 10 100 min
- Particle fluence increases by many orders of magnitude
- Astronauts directly exposed risk lethal radiation doses
- Spacecraft design must include "storm shelters"
- A network of satellites and ground stations monitor the sun for signs of SPEs

## **Violent Space Weather**

Coronal mass ejection, or solar particle event

Image removed due to copyright restrictions.

Figure 4.1 in [SSB-Space Station].

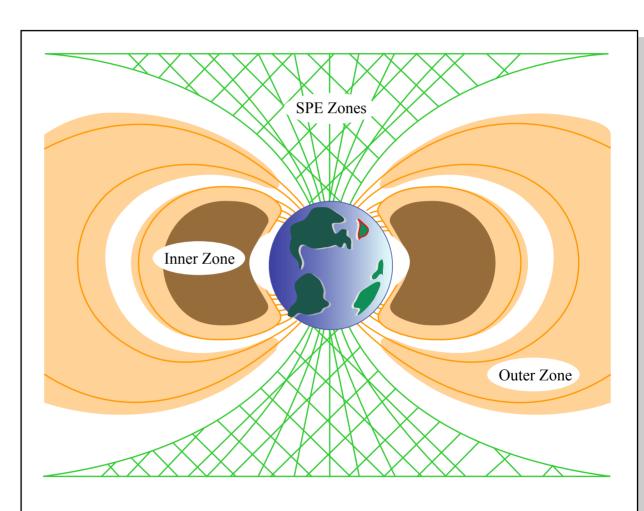
Commission on Physical Sciences, Mathematics, and Applications, Space Studies Board (SSB). *Radiation and the International Space Station: Recommendations to Reduce Risk.* Washington, DC: National Academies Press, 2000.

See http://books.nap.edu/books/0309068851/html/40.html#page\_middle.

Inner Belt mostly protons E ~ 10 MeV

Outer Belt mostly electrons E up to 10 MeV

"Horns" dip in at the poles



The radiation environments of the International Space Station. The figure shows the three regions of space around Earth where penetrating radiation occurs. The inner and outer radiation belts each have an electron and ion component.

Figure by MIT OCW.

Image removed due to copyright restrictions.

- Degrade satellite components
- Background noise in detectors
- Errors in digital circuits
- Electrostatic charge-up in insulators
- Present a threat to astronauts
- •Apollo missions: largest dose component was from travel through the van Allen radiation belts

NASA limits the time spent in high-dose regions of the inner radiation belt.

Image removed due to copyright restrictions.

## The South Atlantic Anomaly

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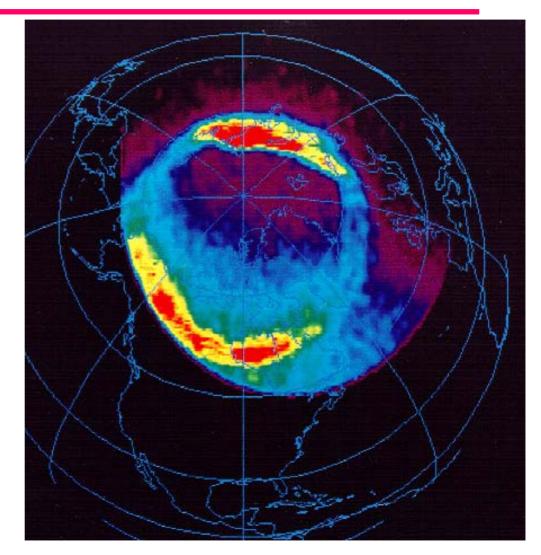
Figure 1.2 in [SSB-Space Station].

See http://books.nap.edu/books/0309068851/html/9.html#pagetop.

#### The Aurora

Image removed due to copyright restrictions.

Energetic charged particles entering the upper atmosphere (~70 miles up) ionize neutral gas molecules.



Courtesy of NASA. "Space Science Photos: Prior to 1997 [cited 29 March 2004] http://www.gsfc.nasa.gov/indepth/photos spaceearly.html

## The International Space Station

# Radiation exposure:

- Radiation Belts
- GCR
- SPEs

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Figure 1.4 in [SSB-Space Station].

See http://books.nap.edu/books/0309068851/html/13.html#page\_bottom.

## The International Space Station

ISS orbit enters higher dose region at high latitudes

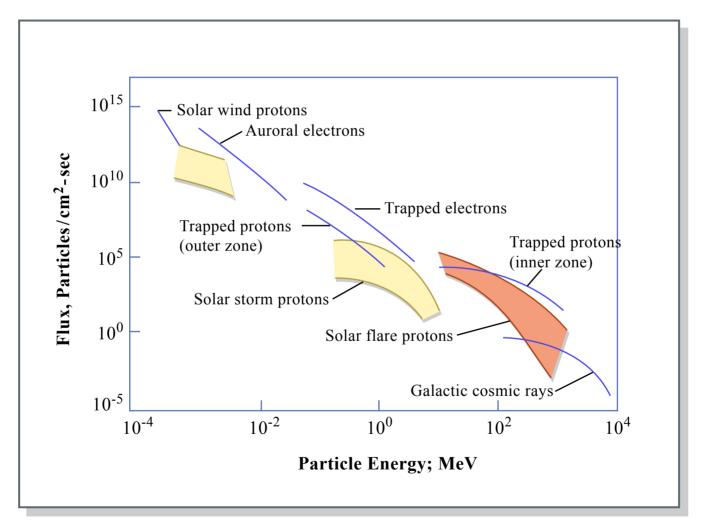
Image removed due to copyright restrictions.

# The International Space Station

Image removed due to copyright restrictions. Figure 1.5 in [SSB-Space Station].

See http://books.nap.edu/books/0309068851/html/14.html#page\_top.Im

## Radiation in Space



## **Biological Effects**

NASA has invested much research effort into the biological effects of the radiations in space.

#### **Protons:**

RBE close to 1 Biological effects fairly well known

#### GCRs:

Fluence is low Biological effects are poorly understood May represent the greatest risk

Fluence rate, outside the earths magnetic field

- 4 protons/cm<sup>2</sup>/sec
- 0.4 helium ions/cm<sup>2</sup>/sec
- 0.04 HZE particles/cm<sup>2</sup>/sec

For a 100  $\mu m^2$  nucleus, every cell nucleus in the body would be hit by:

- a proton once every 3 days
- a helium ion once every month
- an HZE particle once per year

### **NASA's Dilemma**

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# **Shielding**

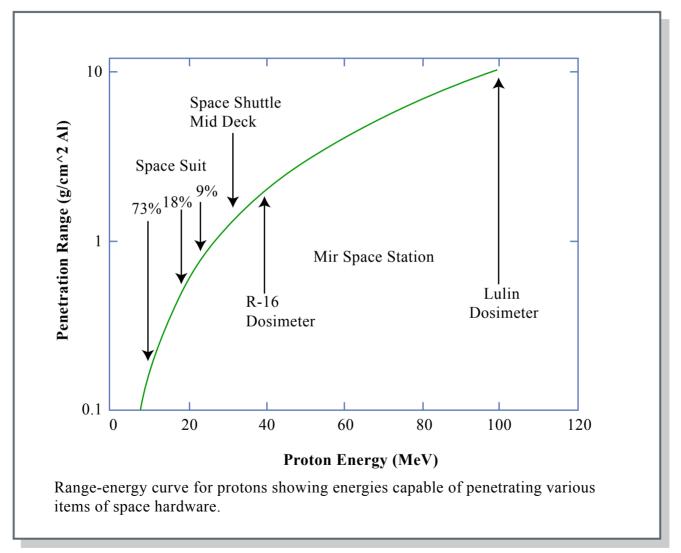


Figure by MIT OCW.

#### The risk to astronauts

Astronauts inside a spacecraft are shielded

The risk is to astronauts outside the spacecraft, or on the surface of the Moon or Mars

Image removed due to copyright restrictions. Figure 2.3 in [SSB-Crew Hazards]. See http://books.nap.edu/books/0309056985/html/16.html#pagetop.

## **Early Biological Effects of Radiation**

#### **Radiation Sickness**

- Occurs within a few hours
- Nausea, vomiting
- Doses: > 1 Sv in less than 1 day

#### **Acute Radiation Syndrome**

- Occurs within 2-4 weeks
- Bone marrow suppression doses: 1.5-2.0 Sv
- Lethal doses (whole-body) 10% at 3 Sv; 90% at 4 Sv (with no countermeasures

#### Skin

- Erythema (reddening) occurs at about 6 Gy
- 15-20 Gy will cause moist desquamation

#### Hair loss

• At doses of ~ 6 Gy or higher

Early effects are unlikely unless an astronaut is exposed while in a non-shielded environment.

### Late Biological Effects of Radiation

Cancer

**CNS** damage

Cataracts: threshold 1.5-2 Gy low-LET (protons similar to low-LET, data in primates)

Late effects are the major concern following exposure to radiation during spaceflights

### Damage to the CNS

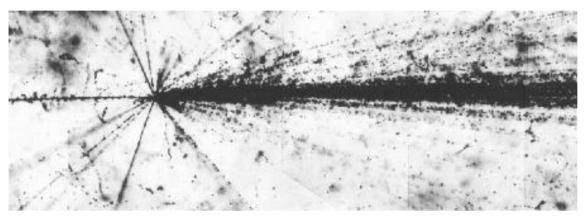
HZE effects in the CNS

Premature aging?

**Experiments underway in rats** 

Image removed due to copyright restrictions.
Figure 2.7 in [SSB-Crew Hazards].
See http://books.nap.edu/books/0309056985/html/25.html#page\_top.Im

# **Shielding**



Source: NASA. "Cosmic Rays." [updated 25 Nov 2001, cited 29 March 2004] http://www-istp.gsfc.nasa.gov/Education/wcosray.html

Track structure of a cosmic ray collision in a nuclear emulsion.

Image removed due to copyright restrictions.

## **Shielding**

GCRs present the greatest shielding problem

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### **Shielding of GCRs**

Shielding ...can make matters worse!

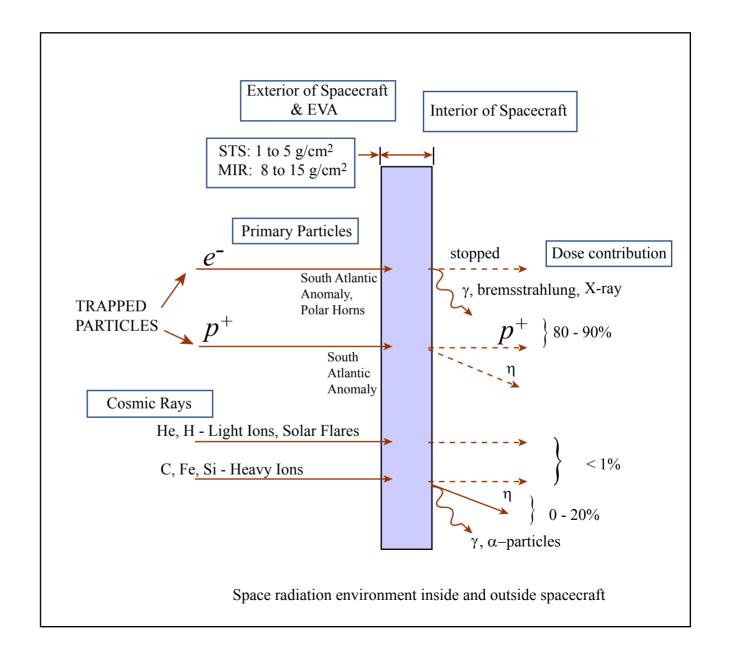


Figure by MIT OCW.

# **Shielding of GCRs**

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Figure 3.1 in [SSB-Crew Hazards].

See http://books.nap.edu/books/0309056985/html/37.html#page\_top.

Shielding for SPEs must be part of the design

Image removed due to copyright restrictions. "Mars TransHab vehicle concept"

#### Characteristics of the inflatable wall and pantry shielding

Inflatable layup and pantry shield materials list				
Inflatable layup	Pantry shield			
Nomex	Acrylic box			
2 layers combitherm/bleeder cloth	Cashews			
Combitherm	Chicken noodle soup			
RTV 560 sealant	Chicken rice soup			
5 layers kevlar	Grilled chicken			
RTV 560 sealant	Peaches			
3 layers 4" polyurethane foam/RTV/Nextel	Fettucine alfredo			
Single aluminized kapton	Broccoli au gratin			
10 layers double aluminized mylar	Strawberries			
Single aluminized kapton	Shortbread cookies			
Beta cloth	Trail mix			

Image removed due to copyright restrictions.

Characteristics of the inflatable wall and pantry shielding. "Mars TransHab vehicle"

### **Exposure limits for Astronauts**

NCRP (1989) Limits for organ dose equivalents (Sv) for low earth orbit exposures

	Blood Forming Organs	Skin	Lens of the eye
Career	1-4	6	4
Annual	0.5	3	2
30 days	0.25	1.5	1

[Average annual background dose to general population is 0.0036 Sv (360 mrem).]

Summary of dose equivalent (cSv) estimates for TransHab inflatable concept

Source	1 y Transit dose equivalent		1.5 y Surface dose equivalent	
	Skin	BFO	Skin	BFO
GCR Solar maximum GCR Solar minimum August 1972 SEP	33.4	27.9	20.1	17.6
	93.8	72.7	46.5	40.7
	63.8	17.0	4.6ª	2.4ª

<sup>&</sup>lt;sup>a</sup> 0-cm and 5-cm depth dose, all others are for CAM.

### **Exposures to Astronauts**

Whole-body dose equivalent (mSv) measured by shuttle passive dosimetry in low-inclination orbits.

Mission	Duration (d)	Low-LET	Neutron	High-LET	Total
STS-4	7.04	0.446	0.156	0.077	0.679
STS-5	5.0	0.278	0.117	0.145	0.540
STS-6	5.0	0.273	0.084	0.138	0.495
STS-7	5.96	0.348	0.014	0.117	0.479
STS-8	6.04	0.348	0.026	0.192	0.566
⟨STS⟩	5.81	0.339	0.079	0.134	0.552

[Annual background is ~3.6 mSv]

### **Exposures to Astronauts**

**Apollo:** 5-12 days 160-1140 mrad 0.01 Sv

Skylab: 20-90 d 1.6-7.7 rad 0.08 Sv

**Shuttle:** 3.2 cGy 0.003 Sv/d

(highest, Hubble repair)

MIR (144 - 468 uGy/day) 0.0005 Sv/d

ISS  $(\sim 0.5 - 2.5 \text{ mGy/day}$  0.0025 Sv/d

Mars mission bone marrow 60-130 cSv\*

\*Exceeds the LEO limit for bone marrow of 50 cSv/y

Average annual background dose

0.0036 Sv

## Summary

The radiation risks to astronauts are real.

NASA is currently investing in radiation biology experiments using high energy Fe ions.

[High dose rate as opposed to the low dose rate in space.]

The biological effects of GCRs are still largely unknown.